

CLUSTERING COMPACT-BINARY OBJECTS IN THE PARAMETER SPACE THROUGH PROBABILISTIC HOUGH TRANSFORM

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Motivation

Masses and Spins distributions of simulated compact-binaries.

$N = 1000$ simulated binaries via population synthesis code, corrupted by realistic noise. These observations are mock posteriors on future LIGO data sets.

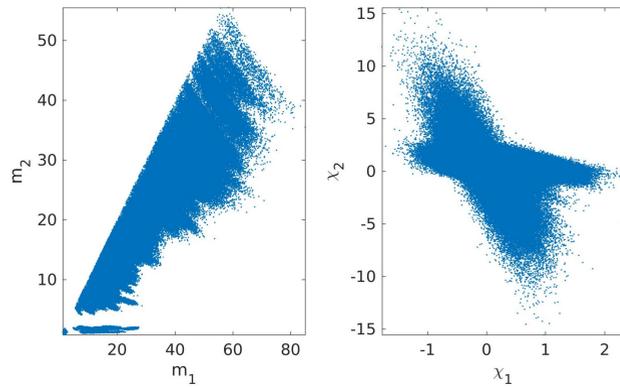


FIGURE 1: Posteriors on future LIGO datasets in masses (left) and spins (right) spaces.

Finding prototypical systems in the parameter space and clustering the models w.r.t. the prototypes could provide useful model-independent information about the population [2].

Building a probabilistic model for each simulated binary.

Each model is a mixture of five gaussians with means aligned along the principal direction and full covariance matrices.

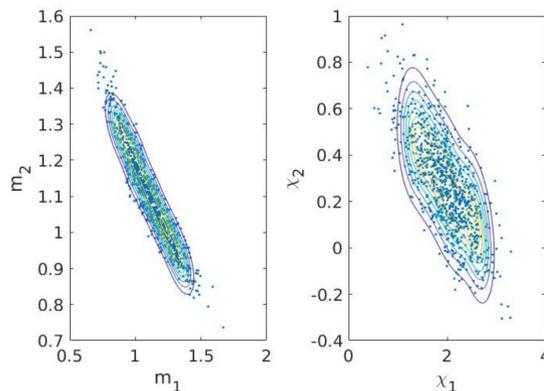


FIGURE 2: Gaussian mixture contours for a single posterior in both parameter spaces.

For each observation we then have a probabilistic model of the form:

$$q_i = \sum_{k=1}^5 p_k \mathcal{N}(\mu_k, \Sigma_k) \quad (1)$$

References

[1] R. T. Ibrahim, P. Tino, R. J. Pearson, T. J. Ponman, and A. Babul. Automated detection of galaxy groups through probabilistic hough transform. In S. Arik, T. Huang, W. K. Lai, and Q. Liu, editors, *Neural Information Processing*, pages 323–331, Cham, 2015. Springer International Publishing.

Methodology

Probabilistic Hough-Transform

Following the work by [1], each parameter space is covered by a uniform grid and each cell (identified by the vector \vec{x}) contains the summation of the responsibilities of all the 1000 models $\mathcal{O}_{i=1}^N$ for that cell:

$$P(\vec{x}) = \sum_{i=1}^N p(i)q(\vec{x}|\mathcal{O}_i) \quad (2)$$

The prior $p(i)$ is assumed to have a flat distribution and is fixed to $1/N$ for every \mathcal{O}_i .

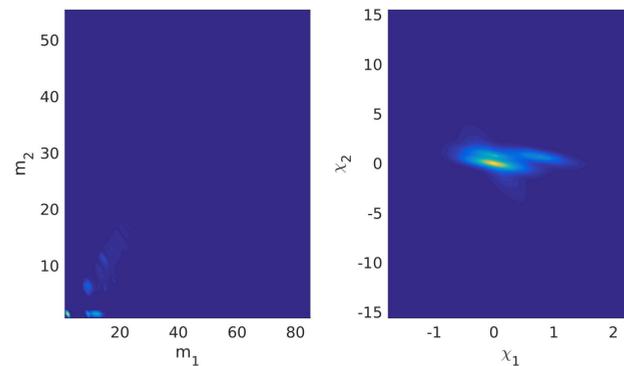


FIGURE 3: PHT for both parameter spaces: masses (left) and spins (right).

The obtained map of the responsibilities on each parameter space is the Probabilistic Hough-Transform (PHT, [3]).

Peak detection and optimal number of clusters

Treating the PHT as a grey-scale image, the number of connected components in it, are the number of peaks (c_j) in the map and thus the clusters prototypes. We can estimate the number of relevant peaks by letting a threshold T vary from the maximum to the minimum of the PHT and looking for a 'knee' in the $T - N_{Peaks}$ curve.

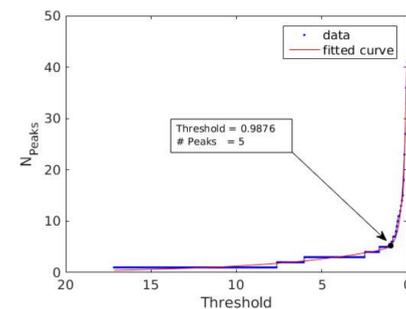


FIGURE 4: $T - N_{Peaks}$ curve for $m_1 - m_2$ parameter space, showing a knee at Threshold ≈ 1 , corresponding to 5 detected peaks.

PHT Peaks and associated clusters

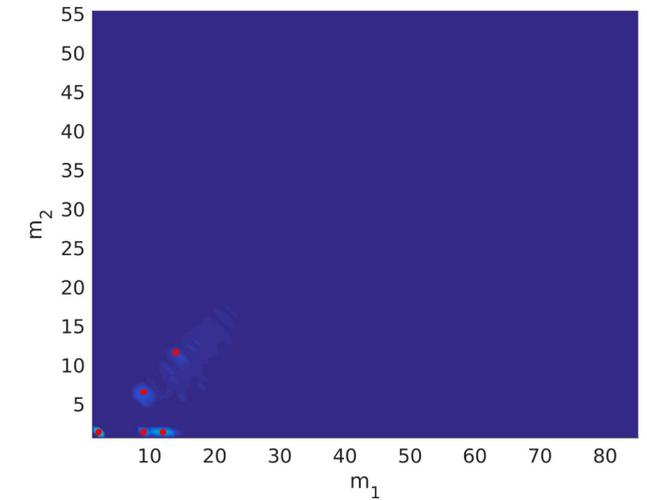


FIGURE 5: PHT with overposed detected optimal peaks.

For each peak we can now compute its responsibility w.r.t. every model:

$$p(c_j|\mathcal{O}_i) = \frac{q(c_j|\mathcal{O}_i)}{\sum_{k=1}^M q(c_k|\mathcal{O}_i)} \quad (3)$$

Results

The clusters are formed by assigning the models to the peak c_j for which $p(c_j|\mathcal{O}_i)$ is maximum, obtaining:

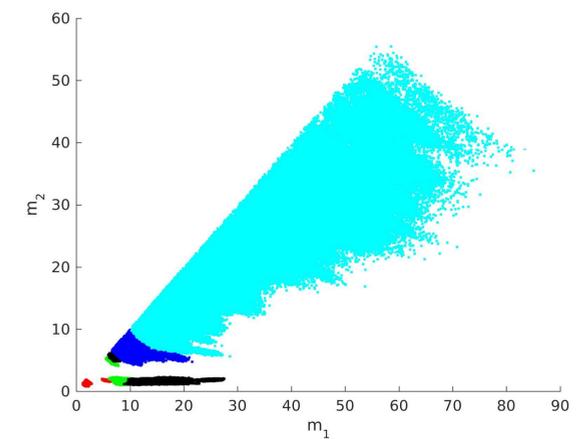


FIGURE 6: Posteriors clustered w.r.t. the detected peaks.

[2] I. Mandel, W. M. Farr, A. Colonna, S. Stevenson, P. Tiño, and J. Veitch. Model-independent inference on compact-binary observations. *MNRAS*, 465(3):3254–3260, Mar 2017.

[3] R. Stephens. Probabilistic approach to the hough transform. *Image and Vision Computing*, 9(1):66–71, 1991. The first BMVC 1990.