

Astro / EPS C12 Midterm I Review Sheet

1. Scale of the cosmos

Solar system - sun, planets, etc.

Milky Way Galaxy - 100 billion stars (incl. Sun)

Local Group - Milky Way + neighboring galaxies

Local Supercluster - 1000's of galaxies (incl. Milky Way)

Universe - everything there is: 100 billion + galaxies

mostly empty space:

distance between planets > size of individual planet

" " stars > " " star

galaxies > " " galaxy

Can't see the entire Universe: only out to 14 billion years $\xleftarrow{\text{light}}$ light year: distance light travels in 1 year
(Universe is 14 billion years old, takes 14 billion years for light to reach us) speed of light: $c = 3 \times 10^8 \text{ m/s}$

Study history of the Universe by looking far away

Origin of universe: Big Bang (14 billion years ago)

only H and He at first, other elements formed by stars

2. Everyday astronomy

Celestial sphere

Local sky: visible half of sphere

Zenith: straight up (overhead)

Horizon: boundary of sky & earth

Earth spins around axis once in 1 day

Seasons

Result of tilt of Earth's axis

most sunlight: summer solstice
(6/21 in northern hemisphere)

least sunlight: winter solstice
(12/21 in north)

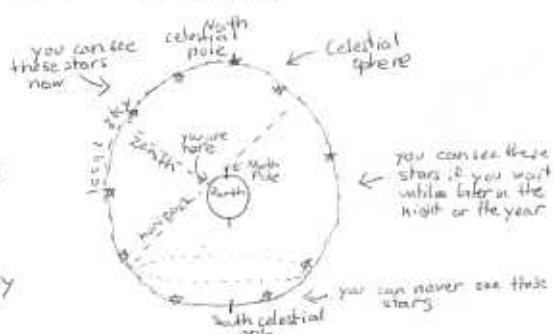
Earth orbits Sun once in 1 year

Moon Phases

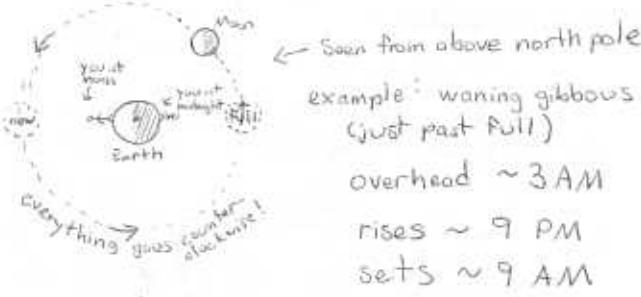
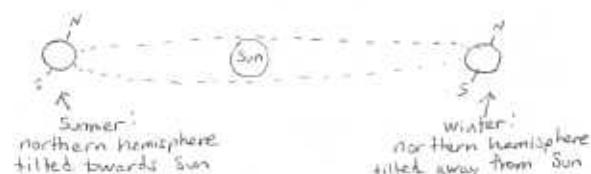
Depends on moon's position relative to the Sun as Moon orbits Earth

Time of moonrise/moonset depends on phase

(Moon is not always up at night!)



$\frac{\text{Sun}}{\text{Earth}} \xrightarrow{1 \text{ AU}}$ $\xrightarrow{\text{Earth}}$ Earth
Astronomical Unit
distance from Earth to Sun
(= 8.3 light-minutes
= 150 million km)

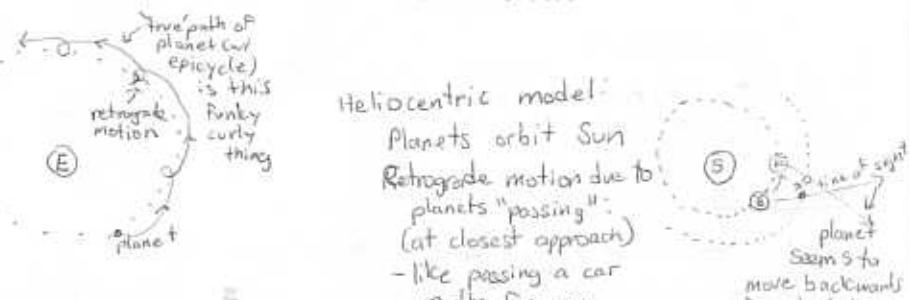


example: waning gibbous
(just past full)

overhead ~3 AM

rises ~9 PM

sets ~9 AM



Planetary motion

Geocentric model

Sun/planets orbit Earth

Ptolemy: planets have "epicycles"
(a sort of secondary orbit)
 \rightarrow retrograde motion

Best model until Copernicus/Kepler

Kepler's Laws:

I. Orbit of a planet is an ellipse.
Sun is at a focus.

II. Planet travels faster nearer the Sun.

$$\text{III. } P^2 = k A^3$$

\uparrow constant
 $= 1$ only if P in years and A in A.U.

P: period
A: semi-major axis
(average distance from Sun)

\rightarrow result:
inner planets move really fast
outer " " " slow

Hypothesis vs.
Theory:
see homework

5
6
planet seems to move backwards for a short time

4. Laws of Matter and Motion

motion:

- speed - rate at which an object is moving
- velocity - speed in a certain direction
- acceleration - change in speed or direction
- momentum - mass \times velocity
- force - change in momentum
- angular momentum - mass \times velocity \times distance from spin axis

Newton's laws of motion:

- Object moves at constant velocity if no net force acts on it (inertia)
- Force = mass \times acceleration
- For any force, there is an equal & opposite reaction force

5. Light

Light is a wave.



$$c = \lambda f$$

- c: speed of light
- λ: wavelength (distance between crests)
- f: frequency (# crests passing per second)

EM spectrum:

Categorize light according to frequency:

radio	low frequency (long λ)	divides visible into colors:
infrared	low frequency (long λ)	red: lower frequency (longer λ)
visible	medium frequency (med λ)	orange yellow green
UV	high frequency (short λ)	blue: higher frequency (shorter λ)
x-ray		violet
gamma ray		

Power:

rate energy is absorbed or emitted

$$E = hf \quad h = 6.6 \times 10^{-34} \frac{\text{J}}{\text{s}}$$

Energy varies with frequency:

Doppler shift

Change in λ and f due to motion along line of sight (towards or away)

object approaches: λ shorter
F higher
"blueshift"

$$\frac{\Delta\lambda}{\lambda_0} = \frac{v_{\text{along line of sight}}}{c}$$

object retreats: λ longer
F lower
"redshift"

8. Origin of Solar System

"Solar nebula theory": collapse of cloud of gas

initially: giant cloud of gas

↓ contracts due to gravity

protoplanetary disk

heating \rightarrow matter falls inward: potential (conservation of energy) thermal

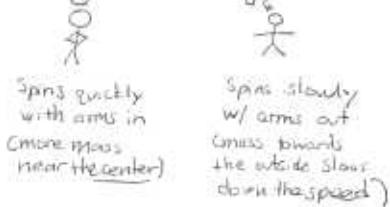
spinning \rightarrow matter falls inward: conservation of angular momentum says it must spin faster

flattening \rightarrow collisions force all orbits to be in same plane and nearly circular orbits

↓ result:
no flat, spinning disk that is hot in the center

↓ Sun forms in center, planets in the disk

Conservation of angular momentum

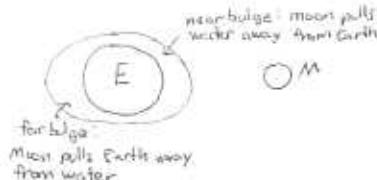


types of energy

- kinetic - energy of motion
- radiative - energy of light
- potential - stored energy
- thermal - energy of heat

Tides:

gravitational pull of Moon and Sun on oceans



highest tide when Sun's pull and moon's pull work together (new moon)

Matter can:

- absorb light (absorb energy)
- emit light (release energy)
- reflect light (bounce off surface)
- scatter in air/water

Spectrum:

plot of wavelength vs. amount of light observed with that wavelength



Important types of spectra:

blackbody - continuous

"bump": size and position of bump (i.e., wavelength and total amount of energy) depend on temperature.

higher T \rightarrow more energy emitted
peak move to shorter λ
normally in infrared part of spectrum

Emission lines - "spikes" or visually colorful lines against background, due to electron jumping from high energy level to lower one:



The pattern/wavelengths of lines tells you the composition of the gas that's doing the emitting/describing, like a "fingerprint" of the atom

Absorption lines - downward spikes where light is missing from a continuous spectrum. Due to photons hitting atoms to boost an electron to a higher level!



"First line":

most heating is in center (Sun is burning) \rightarrow temperatures in the inner disk are high: cannot form ice only rocky objects can solidify. Cannot retain hydrogen/helium in outer disk, H₂O can condense as ice \rightarrow much faster accretion of H/He.

\rightarrow rocky planets closer to Sun
gaseous planets & icy moons further