

# Bibliography

- Ade, P., et al., First Season QUaD CMB Temperature and Polarization Power Spectra, *apj*, 674, 22–28, doi:10.1086/524922, 2008.
- Agilent, *Advanced Design Suite Momentum v6u2, User Manual*, Agilinet, 2006.
- Apostle, T., *Calculus, Vol. 1: One-Variable Calculus with an Introduction to Linear Algebra*, Wiley, 1967.
- Arnold, K., personal communication, 2008.
- Behdad, N., and K. Sarabandi, Wideband double-element ring slot antenna, *Electronics Letters*, 40(7), 408 – 409, doi:10.1049/el:20040292, 2004.
- Bell, R., *Introductory Fourier Transform Spectroscopy*, Academic Pr, 1972.
- Benford, D. J., M. C. Gaidis, and J. W. Kooi, Optical properties of Zitex in the infrared to submillimeter, *ao*, 42, 5118–5122, doi:10.1364/AO.42.005118, 2003.
- Benson, B. A., et al., Peculiar Velocity Limits from Measurements of the Spectrum of the Sunyaev-Zeldovich Effect in Six Clusters of Galaxies, *apj*, 592, 674–691, doi:10.1086/375864, 2003.
- Bond, J., I could tell you, but then I'd have to kill you, 2010.
- Booker, H., Slot aerials and their relation to complementary wire aerials (Babinet's principle), *Electrical Engineers - Part IIIA: Radiolocation, Journal of the Institution of*, 93(4), 620 –626, doi:10.1049/ji-3a-1.1946.0150, 1946.
- Born, M., and E. Wolf, *Principles of Optics: Electromagnetic Theory of Propagation, Interference and Diffraction of Light*, Cambridge University Press, 1999.
- Bussmann, R. S., W. L. Holzapfel, and C. L. Kuo, Millimeter Wavelength Brightness Fluctuations of the Atmosphere above the South Pole, *apj*, 622, 1343–1355, doi:10.1086/427935, 2005.
- Carlstrom, J. E., et al., The 10 Meter South Pole Telescope, *ArXiv e-prints*, 2009.
- Carroll, S., *Spacetime and Geometry: An Introduction to General Relativity*, Benjamin Cummings, 2003.
- Challinor, A., and H. Peiris, Lecture notes on the physics of cosmic microwave background anisotropies, in *American Institute of Physics Conference Series, American Institute of Physics Conference Series*, vol. 1132, edited by M. Novello & S. Perez, pp. 86–140, doi: 10.1063/1.3151849, 2009.

- Chang, J., Xenon Difluoride Etching System (XETCH), *Manual Nanoolab Manual section 7.13*, UC Berkeley, Berkeley, Ca, USA, 1998.
- Chang, J., CPA Sputtering System, *Manual Nanoolab Manual section 6.04*, UC Berkeley, Berkeley, Ca, USA, 2010a.
- Chang, J., Oxford Plasmalab 80plus PECVD System, *Manual Nanoolab Manual section 6.29*, UC Berkeley, Berkeley, Ca, USA, 2010b.
- Chang, J., Tystar17 Non-MOS Low Stress Nitride and High Temp. Oxide LPCVD Furnace, *Manual Nanoolab Manual section 5.17*, UC Berkeley, Berkeley, Ca, USA, 2010c.
- Chattopadhyay, G., and J. Zmuidzinas, A dual-polarized slot antenna for millimeter waves, *Antennas and Propagation, IEEE Transactions on*, 46(5), 736 –737, doi:10.1109/8.668920, 1998.
- Chattopadhyay, G., J. Glenn, J. Bock, B. Rownd, M. Caldwell, and M. Griffin, Feed horn coupled bolometer arrays for SPIRE - design, simulations, and measurements, *Microwave Theory and Techniques, IEEE Transactions on*, 51(10), 2139 – 2146, doi:10.1109/TMTT.2003.817428, 2003.
- Chattopadhyay, G., C.-L. Kuo, P. Day, J. Bock, J. Zmuidzinas, and A. Lange, Planar antenna arrays for CMB polarization detection, in *Infrared and Millimeter Waves, 2007 and the 2007 15th International Conference on Terahertz Electronics. IRMMW-THz. Joint 32nd International Conference on*, pp. 184 –185, 2007.
- Chervenak, J. A., K. D. Irwin, E. N. Grossman, J. M. Martinis, C. D. Reintsema, and M. E. Huber, Superconducting multiplexer for arrays of transition edge sensors, *Applied Physics Letters*, 74(26), 4043–4045, doi:10.1063/1.123255, 1999.
- Chiang, H. C., et al., Measurement of Cosmic Microwave Background Polarization Power Spectra from Two Years of BICEP Data, *apj*, 711, 1123–1140, doi:10.1088/0004-637X/711/2/1123, 2010.
- Collin, R., *Antennas and Radiowave Propagation*, Mcgraw-Hill, 1985.
- de Bernardis, P., et al., Detection of anisotropy in the Cosmic Microwave Background at horizon and sub-horizon scales with the BOOMERanG experiment, *ArXiv Astrophysics e-prints*, 2000.
- Deschamps, G., Impedance properties of complementary multiterminal planar structures, *Antennas and Propagation, IRE Transactions on*, 7(5), 371 –378, doi:10.1109/TAP.1959.1144717, 1959.
- Dodelson, S., *Modern cosmology*, Academic Press, 2003.
- Doe, J., <http://universe-review.ca/>, 2010.
- DuHamel, R., Dual polarized sinuous antennas, United States Patent 4658262, 1987.
- DuHamel, R., and D. Isbell, Broadband logarithmically periodic antenna structures, in *IRE International Convention Record*, vol. 5, pp. 119 – 128, 1957.

- Dyson, J., The equiangular spiral antenna, *Antennas and Propagation, IRE Transactions on*, 7(2), 181–187, doi:10.1109/TAP.1959.1144653, 1959.
- Edwards, J., personal communication, 2008.
- Engargiola, G., Tapered microstrip balun for integrating a low noise amplifier with a nonplanar log periodic antenna, *Review of Scientific Instruments*, 74, 5197–5200, doi:10.1063/1.1622975, 2003.
- Engargiola, G., W. Holzapfel, A. Lee, M. Myers, R. O’Brien, P. Richards, H. Tran, and H. Spieler, Planar channelized log-periodic antenna, in *Microwave, Antenna, Propagation and EMC Technologies for Wireless Communications, 2005. MAPE 2005. IEEE International Symposium on*, vol. 1, pp. 306–309 Vol. 1, doi:10.1109/MAPE.2005.1617910, 2005.
- Filipovic, D., S. Gearhart, and G. Rebeiz, Double-slot antennas on extended hemispherical and elliptical silicon dielectric lenses, *Microwave Theory and Techniques, IEEE Transactions on*, 41(10), 1738–1749, doi:10.1109/22.247919, 1993.
- Finelli, F., J. Hamann, S. M. Leach, and J. Lesgourgues, Single-field inflation constraints from CMB and SDSS data, *Journal of Cosmology and Astro-Particle Physics*, 4, 11–+, doi:10.1088/1475-7516/2010/04/011, 2010.
- Finkbeiner, D. P., M. Davis, and D. J. Schlegel, Extrapolation of Galactic Dust Emission at 100 Microns to Cosmic Microwave Background Radiation Frequencies Using FIRAS, *apj*, 524, 867–886, doi:10.1086/307852, 1999.
- Fixsen, D. J., E. S. Cheng, J. M. Gales, J. C. Mather, R. A. Shafer, and E. L. Wright, The Cosmic Microwave Background Spectrum from the Full COBE FIRAS Data Set, *apj*, 473, 576–+, doi:10.1086/178173, 1996.
- Galbraith, C., Cochlear-inspired channelizing filters for Wideband Radio Systems, Ph.D. thesis, The University Michigan Ann Arbor, 2008.
- Galbraith, C., and G. Rebeiz, Higher Order Cochlea-Like Channelizing Filters, *Microwave Theory and Techniques, IEEE Transactions on*, 56(7), 1675–1683, doi:10.1109/TMTT.2008.925574, 2008.
- Galbraith, C., R. White, L. Cheng, K. Grosh, and G. Rebeiz, Cochlea-Based RF Channelizing Filters, *Circuits and Systems I: Regular Papers, IEEE Transactions on*, 55(4), 969–979, doi:10.1109/TCSI.2008.916537, 2008.
- Gitin, M., F. Wise, G. Arjavalingham, Y. Pastol, and R. Compton, Broad-band characterization of millimeter-wave log-periodic antennas by photoconductive sampling, *Antennas and Propagation, IEEE Transactions on*, 42(3), 335–339, doi:10.1109/8.280719, 1994.
- Goldin, A., J. J. Bock, C. Hunt, A. E. Lange, H. Leduc, A. Vayonakis, and J. Zmuidzinas, SAMBA: Superconducting antenna-coupled, multi-frequency, bolometric array, *Low Temperature Detectors*, 605, 251–254, doi:10.1063/1.1457640, 2002.
- Goodman, J., *Introduction to Fourier Optics*, McGraw-Hill Science, 1968.

- Griffin, M. J., J. J. Bock, and W. K. Gear, Relative Performance of Filled and Feedhorn-Coupled Focal-Plane Architectures, *Appl. Opt.*, *41*(31), 6543–6554, 2002.
- Guildemeister, J., Voltage-Biased Superconducting Bolometers for infrared and mm-waves, Ph.D. thesis, University of California Berkeley, 2000.
- Gupta, K., G. Garg, B. Bahl, and P. Bhartia, *Microstrip Lines and Slotlines 2nd Ed.*, Artech House Publishers, 1996.
- Halverson, N., NEP and Mapping speed for 92 GHz and 150 GHz, *Internal Technical Memo EPEX-SZ-041211a*, UC Berkeley, Berkeley, Ca, USA, 2004.
- Halverson, N. W., et al., Degree Angular Scale Interferometer First Results: A Measurement of the Cosmic Microwave Background Angular Power Spectrum, *apj*, *568*, 38–45, doi:10.1086/338879, 2002.
- Hamilton, B., Plasma-Therm Parallel Plate Etcher, *Manual Nanoolab Manual section 7.09*, UC Berkeley, Berkeley, Ca, USA, 2010.
- Hanany, S., et al., MAXIMA-1: A Measurement of the Cosmic Microwave Background Anisotropy on Angular Scales of  $10^{\circ}$ -5deg, *apjl*, *545*, L5–L9, doi:10.1086/317322, 2000.
- Hu, W., and S. Dodelson, Cosmic Microwave Background Anisotropies, *araa*, *40*, 171–216, doi:10.1146/annurev.astro.40.060401.093926, 2002.
- Hu, W., and M. White, A CMB polarization primer, *New Astronomy*, *2*, 323–344, doi:10.1016/S1384-1076(97)00022-5, 1997.
- Irwin, K., and G. Hilton, Transition-Edge Sensors, in *Cryogenic Particle Detection*, pp. 63–149, Springer-Verlag, Heidelberg, 2005.
- Jackson, J. D., *Classical Electrodynamics Third Edition*, Wiley, 1998.
- Johnson, B. R., et al., MAXIPOL: a balloon-borne experiment for measuring the polarization anisotropy of the cosmic microwave background radiation, *New Astronomy Reviews*, *47*(11-12), 1067 – 1075, doi:DOI:10.1016/j.newar.2003.09.034, 2003.
- Kaplinghat, M., L. Knox, and Y. Song, Determining Neutrino Mass from the Cosmic Microwave Background Alone, *Physical Review Letters*, *91*(24), 241,301–+, doi:10.1103/PhysRevLett.91.241301, 2003.
- Kerr, A., Surface Impedance of Superconductors and Normal Conductors in EM Simulators, *Internal Technical Memo Alma Memo No. 245*, National Radio Astronomy Observatory, Green Bank, W. Va, USA, 1999.
- Kittel, C., *Introduction to Solid State Physics, 8th Ed*, wiley, 2004.
- Kogut, A., et al., PAPP: Primordial anisotropy polarization pathfinder array, *New Astronomy Review*, *50*, 1009–1014, doi:10.1016/j.newar.2006.09.024, 2006.
- Kogut, A., et al., Three-Year Wilkinson Microwave Anisotropy Probe (WMAP) Observations: Foreground Polarization, *apj*, *665*, 355–362, doi:10.1086/519754, 2007.

- Komatsu, E., et al., Seven-Year Wilkinson Microwave Anisotropy Probe (WMAP) Observations: Cosmological Interpretation, *ArXiv e-prints*, 2010.
- Kormanyos, B., P. Ostdiek, W. Bishop, T. Crowe, and G. Rebeiz, A planar wideband 80-200 GHz subharmonic receiver, *Microwave Theory and Techniques, IEEE Transactions on*, *41*(10), 1730–1737, doi:10.1109/22.247918, 1993.
- Kosowsky, A., Cosmic microwave background polarization., *Annals of Physics*, *246*, 49–85, doi:10.1006/aphy.1996.0020, 1996.
- Kovac, J. M., Detection of polarization in the cosmic microwave background using DASI, Ph.D. thesis, THE UNIVERSITY OF CHICAGO, 2004.
- Kovac, J. M., E. M. Leitch, C. Pryke, J. E. Carlstrom, N. W. Halverson, and W. L. Holzapfel, Detection of polarization in the cosmic microwave background using DASI, *nat*, *420*, 72–787, doi:10.1038/nature01269, 2002.
- Kumar, S., Submillimeter wave camera using a novel photon detector technology, Ph.D. thesis, California Institute of Technology, 2007.
- Lanting, T., Multiplexed Readout of Superconducting Bolometers for Cosmological Observation, Ph.D. thesis, University of California Berkeley, 2006.
- Lanting, T. M., et al., Frequency-domain multiplexed readout of transition-edge sensor arrays with a superconducting quantum interference device, *Applied Physics Letters*, *86*(11), 112511, doi:10.1063/1.1884746, 2005.
- Larson, D., et al., Seven-Year Wilkinson Microwave Anisotropy Probe (WMAP) Observations: Power Spectra and WMAP-Derived Parameters, *ArXiv e-prints*, 2010a.
- Larson, D., et al., Seven-Year Wilkinson Microwave Anisotropy Probe (WMAP) Observations: Power Spectra and WMAP-Derived Parameters, *ArXiv e-prints*, 2010b.
- Lee, A. T., et al., POLARBEAR: Ultra-high Energy Physics with Measurements of CMB Polarization, in *American Institute of Physics Conference Series, American Institute of Physics Conference Series*, vol. 1040, edited by H. Kodama & K. Ioka, pp. 66–77, doi: 10.1063/1.2981555, 2008.
- Lewis, A., and A. Challinor, Weak gravitational lensing of the CMB, *physrep*, *429*, 1–65, doi:10.1016/j.physrep.2006.03.002, 2006.
- Liddle, A. R., An Introduction to Cosmological Inflation, in *High Energy Physics and Cosmology, 1998 Summer School*, edited by R. L. Walsworth & D. F. Phillips, pp. 260–+, 1999.
- Liddle, A. R., and D. H. Lyth, *Cosmological Inflation and Large-Scale Structure*, Cambridge University Press, 2000.
- Lis, D., CSO Atmospheric Transmission Interactive Plotter, <http://www.submm.caltech.edu/cso/weather/atplot.shtml>, 2010.

- Liu, L., H. Xu, R. R. Percy, D. L. Herald, A. W. Lichtenberger, J. L. Hesler, and R. M. Weikle, Development of Integrated Terahertz Broadband Detectors Utilizing Superconducting Hot-Electron Bolometers, *IEEE Transactions on Applied Superconductivity*, 19, 282–286, doi:10.1109/TASC.2009.2018268, 2009.
- Ludwig, A., The definition of cross polarization, *Antennas and Propagation, IEEE Transactions on*, 21(1), 116 – 119, 1973.
- Lyth, D. H., Introduction to Cosmology, *ArXiv Astrophysics e-prints*, 1993.
- Matthaei, G., E. Jones, and L. Young, *Microwave Filters, Impedance-Matching Networks, and Coupling Structures*, Artech House Publishers, 1980.
- McGinnis, D., and J. Beyer, A broad-band microwave superconducting thin-film transformer, *Microwave Theory and Techniques, IEEE Transactions on*, 36(11), 1521 –1525, doi:10.1109/22.8916, 1988.
- Mehl, J., et al., TES Bolometer Array for the APEX-SZ Camera, *Journal of Low Temperature Physics*, 151, 697–702, doi:10.1007/s10909-008-9738-1, 2008.
- Meng, X., GCA 6200 Wafer Stepper, *Manual Nanoolab Manual section 4.13*, UC Berkeley, Berkeley, Ca, USA, 2010.
- Miller, A. D., et al., A Measurement of the Angular Power Spectrum of the Cosmic Microwave Background from  $L = 100$  to 400, *apjl*, 524, L1–L4, doi:10.1086/312293, 1999.
- Mushiake, Y., *Self-Complementary Antennas: Principle of Self-Complementarity for Constant Impedance*, New York Academic Press, 1996.
- Myers, M., personal communication, 2008.
- Myers, M. J., et al., An antenna-coupled bolometer with an integrated microstrip bandpass filter, *Applied Physics Letters*, 86(11), 114103, doi:10.1063/1.1879115, 2005.
- Nurnberger, M., and J. Volakis, A new planar feed for slot spiral antennas, *Antennas and Propagation, IEEE Transactions on*, 44(1), 130 –131, doi:10.1109/8.477538, 1996.
- O’Brient, R., et al., A Multi-Band Dual-Polarized Antenna-Coupled TES Bolometer, *Journal of Low Temperature Physics*, 151, 459–463, doi:10.1007/s10909-007-9698-x, 2008a.
- O’Brient, R., et al., Sinuous antennas for cosmic microwave background polarimetry, in *Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series*, vol. 7020, doi: 10.1117/12.788526, 2008b.
- Olive, K. A., G. Steigman, and T. P. Walker, Primordial nucleosynthesis: theory and observations, *physrep*, 333, 389–407, doi:10.1016/S0370-1573(00)00031-4, 2000.
- Pardo, J., J. Cernicharo, and E. Serabyn, Atmospheric transmission at microwaves (ATM): an improved model for millimeter/submillimeter applications, *Antennas and Propagation, IEEE Transactions on*, 49(12), 1683 –1694, doi:10.1109/8.982447, 2001.
- Partridge, R., *3K: The Cosmic Microwave Background Radiation*, Cambridge University Press, Palo Alto, 1995.

- Pathria, R., *Statistical Mechanics, 2nd Edition*, Butterworth-Heinemann, 1996.
- Pearson, T. J., et al., The Anisotropy of the Microwave Background to  $l = 3500$ : Mosaic Observations with the Cosmic Background Imager, *apj*, 591, 556–574, doi:10.1086/375508, 2003.
- Pisano, G., G. Savini, P. A. R. Ade, V. Haynes, and W. K. Gear, Achromatic half-wave plate for submillimeter instruments in cosmic microwave background astronomy: experimental characterization, *ao*, 45, 6982–6989, doi:10.1364/AO.45.006982, 2006.
- Pozar, D., *Microwave Engineering*, Wiley, 2004.
- Raman, S., and G. Rebeiz, Single- and dual-polarized millimeter-wave slot-ring antennas, *Antennas and Propagation, IEEE Transactions on*, 44(11), 1438–1444, doi:10.1109/8.542067, 1996.
- Rauscher, C., Efficient design methodology for microwave frequency multiplexers using infinite-array prototype circuits, *Microwave Theory and Techniques, IEEE Transactions on*, 42(7), 1337–1346, doi:10.1109/22.299727, 1994.
- Reichardt, C. L., et al., High-Resolution CMB Power Spectrum from the Complete ACBAR Data Set, *apj*, 694, 1200–1219, doi:10.1088/0004-637X/694/2/1200, 2009.
- Richards, P. L., Bolometers for infrared and millimeter waves, *Journal of Applied Physics*, 76(1), 1–24, doi:10.1063/1.357128, 1994.
- Rumsey, V., *Frequency Independent Antennas*, New York Academic Press, 1966.
- Rybicki, G. B., and A. P. Lightman, *Radiative processes in astrophysics*, Wiley, 1979.
- Saini, L., and R. Bradley, The Sinuous Antenna - a Dual Polarized Element for Wideband Phased Array Feed Application, *Internal Technical Memo Electronic Division Internal Report-No. 31*, National Radio Astronomy Observatory, Green Bank, W. Va, USA, 1996.
- Sarpeshkar, R., R. Lyon, and C. Mead, *A Low-Power Wide-Dynamic-Range Analog VLSI Cochlea*, vol. 447, pp. 49–103, Springer, doi:10.1007/b102308, 2007.
- Simons, R., *Coplanar Waveguide Circuits Components and Systems*, Wiley-IEEE Press, 2001.
- Smoot, G. F., M. V. Gorenstein, and R. A. Muller, Detection of anisotropy in the cosmic blackbody radiation, *Physical Review Letters*, 39, 898–901, doi:10.1103/PhysRevLett.39.898, 1977.
- Smoot, G. F., et al., Structure in the COBE differential microwave radiometer first-year maps, *apjl*, 396, L1–L5, doi:10.1086/186504, 1992.
- Staniszewski, Z., et al., Galaxy Clusters Discovered with a Sunyaev-Zel’dovich Effect Survey, *apj*, 701, 32–41, doi:10.1088/0004-637X/701/1/32, 2009.
- Swanson, D., Thin-film lumped-element microwave filters, in *Microwave Symposium Digest, 1989.*, *IEEE MTT-S International*, pp. 671–674 vol.2, doi:10.1109/MWSYM.1989.38814, 1989.

- Takahashi, Y., personal communication, 2008.
- Van Duzer, T., *Principles of Superconducting Devices and Circuits*, Prentice Hall, 1998.
- Vieira, J. D., et al., Extragalactic millimeter-wave sources in South Pole Telescope survey data: source counts, catalog, and statistics for an 87 square-degree field, *ArXiv e-prints*, 2009.
- Watts, L., Cochlear Mechanics, Analysis and Analog VLSI, Ph.D. thesis, California Institute of Technology, 1993.
- Weisend, J., *The Handbook Of Cryogenic Engineering*, CRC Press, 1998.
- Werthamer, N. R., Theory of the Superconducting Transition Temperature and Energy Gap Function of Superposed Metal Films, *Phys. Rev.*, *132*(6), 2440–2445, doi:10.1103/PhysRev.132.2440, 1963.
- Woodcraft, A. L., and A. Gray, A low temperature thermal conductivity database, in *A low temperature thermal conductivity database*, vol. 1185, edited by B. Young, B. Cabrera, and A. Miller, pp. 681–684, AIP, doi:10.1063/1.3292433, 2009.
- Woody, D. P., and P. L. Richards, Spectrum of the cosmic background radiation, *Physical Review Letters*, *42*, 925–929, doi:10.1103/PhysRevLett.42.925, 1979.
- Yassin, G., and S. Withington, Loss In Normal And Superconducting Millimetre-Wave And Submillimetre Wave Microstrip Transmission Line, in *Computation in Electromagnetics, Third International Conference on (Conf. Publ. No. 420)*, pp. 149–154, 1996.
- Yoon, K. W., et al., Feedhorn-Coupled TES Polarimeters for Next-Generation CMB Instruments, in *American Institute of Physics Conference Series, American Institute of Physics Conference Series*, vol. 1185, edited by B. Young, B. Cabrera, & A. Miller, pp. 515–518, doi:10.1063/1.3292392, 2009.
- Zmuidzinas, J., Cramér-Rao sensitivity limits for astronomical instruments: implications for interferometer design, *Journal of the Optical Society of America A*, *20*, 218–233, doi:10.1364/JOSAA.20.000218, 2003a.
- Zmuidzinas, J., Thermal noise and correlations in photon detection, *ao*, *42*, 4989–5008, doi:10.1364/AO.42.004989, 2003b.
- Zmuidzinas, J., personal communication, 2010.



## Appendix A

# MATLAB raytracing software

### A.1 Overview

This appendix contains a hardcopy of the raytracing script and all functions. This code is included here as a reference and to ensure that it is not lost. It was written with Matlab v 7.0. The function progmeter.m is needed as well, but was written by another author and is available for download from <http://www.mathworks.com/matlabcentral/fileexchange>.

### A.2 Main script

```
1 %BEAMSCRIPT
2 %
3 %This script modifies the far-field antenna patterns from ADS-momentum to
4 %account for a contacting extended hemispherical lens. It accounts for
5 %refraction at the lens surface as well as diffraction. The user should
6 %modify the *non-indented* variables at the beginning of the script between
7 %the BEGIN and END USER DEFINED INPUT
8 %
9 %This script uses the following functions which must be present in the same
10 %directory for it to function properly: build_lens.m, raytrace.m,
11 %readfff.m, normalize.m, rect2plane_incidence.m, plane_incidence2rect.m,
12 %rect2sphere.m, sphere2rect.m, refraction_nocoat.m, refraction_lcoat.m,
13 %refraction_3coat.m, matrix_mult4d.m, surfacecurrents.m, Diffraction.m,
14 %progmeter.m, writefff.m
15 %
16 %Because the program was developed on a laptop with 512MB RAM, there was
17 %insufficient memory to store all variables at once, so a series of loops
18 %are used instead, with a timer so the user can watch progress. If all
19 %discretization is done at 3 degrees, then the integral takes 30 sec to
20 %finish.
21 %Special thanks to Jen Edwards for help troubleshooting this code, pointing
22 %out that Matlab's dot function conjugates it's first argument, and
23 %suggesting the use of image currents.
24 %Roger O'Brient          Jan 2010
25
26 home;
```

```

27 clear all;
28
29 c=3*10^8;%[m/s] speed of light free space
30 eta=377;%[Ohms] impedance free space
31
32 f_GHz=150; %[GHz] Frequency
33 f=f_GHz*10^9; %[Hz]
34 f_str=[num2str(round(f_GHz))];
35
36 %Lens Properties
37 er=11.7; % rel permativity of lens material (silicon)
38 %er=10.5;%rel permativity if lens material (sintered alumina)
39     nlens=sqrt(er); %index refraction of lens material
40 R=13.7/2; %[mm] Lens Radius
41     R=R/1000; %[m] Lens Radius
42 %Lext=.2767*R; %[mm] Hyperhemisphere for er=11.7
43 Lext=.3876*R; %[mm] Synth ellipse for er=11.7
44
45 %Change coating-flag for the three posible coatings:
46 %coatingflag='no coating';
47 coatingflag='single layer';
48 %coatingflag='three layers';
49
50 %Single Layer AR-coating properities
51 thickAR=0.30/2; %[mm] Currently lambda/4 for 300GHz
52     thickAR=thickAR/1000; %[m]
53 nAR=2; %Index of stycast-2850, Lamb compendium data, 4.8K, 100GHz
54
55 %Multi-layer AR properities. I've set all thicknesses to lambda/4 for
56 %160GHz center frequency
57 thick1=0.19; %[mm] TMM6
58 thick2=0.27; %[mm] TMM3
59 thick3=0.3906; %[mm] Zitex
60     thick1=thick1/1000; %[m]
61     thick2=thick2/1000; %[m]
62     thick3=thick3/1000; %[m]
63 n1=2.45; %index refraction of TMM6, from Erin Quealy
64 n2=1.73; %index refraction of TMM3, from Erin Quealy
65 n3=1.2; %%index refraction of Zitex, from Dominec Benford et all.
66
67 %antenna location relative to hemispherical center
68 centerx=0; %[mm]
69 centery=0; %[mm]
70 centerz=-Lext; %[mm]
71     ant_loc=[centerx;centery;centerz] ;
72
73 %To include image currents, set inc_img=1. To exclude, set it to 0
74 inc_img=1 ;
75
76 %Angular steps for integration and far field pattern display
77 %Code takes 30-40 sec with all set to 3 deg.
78 hemthetastep=3; %[deg]
79 hemphistep=3; %[deg]
80     hemthetastep=pi/180*hemthetastep; %[rad]
81     hemphistep=pi/180*hemphistep; %[rad]
82 % PATTERN THETA/PHI
83 ffthetastep=3; %[deg]
84 ffphistep=3; %[deg]
85     ffthetastep=pi/180*ffthetastep; %[rad]

```

```

86     ffphistep=pi/180*ffphistep; %[rad]
87
88 %Input file from ADS momentum
89 pathin='TestResults\';
90 fnamein=[ 'Momentum_input.fff' ] ;
91
92 %Output file to write to.
93 pathout=pathin;
94 fnameout=['Sample_output.', f_str, 'GHz.txt'];
95
96 commentary = { [ 'Frequency = ' num2str(f.GHz), ' GHz' ];
97                [ 'er = ', num2str(er,'%2.2f') ] ;
98                [ 'Lens Radius = ' , num2str(R*1e3,'%2.4f'), ' mm, ' ,...
99                  'Extension Length = ' , num2str(Lext*1e3,'%2.4f'),' mm '];
100               [ 'antenna at x = ', num2str(centerx*1e3,'%2.6f'),' mm, '...
101                 ' y = ', num2str(centery*1e3,'%2.6f'),...
102                 ' mm relative to lens central axis' ] ;
103               [ 'Antireflection model: ', coatingflag]; } ;
104
105
106
107 disp('Performing Raytracing inside the lens') ;
108 tic ;
109
110 %Diffraction happens at the outer-most surface, so we need to adjust the
111 %radius of the lens surface according to the number of layers in the
112 %AR-coating.
113 switch coatingflag
114     case {'no coating'}
115         Rsurf=R;
116     case {'single layer'}
117         Rsurf=R+thickAR;
118     case {'three layers'}
119         Rsurf=R+thick1+thick2+thick3;
120 end
121 %Calculate lens geometry:
122 %hemtheta,hemphi: angular postions of surface patches to hemispherical
123 %    center
124 %normal=unit vector normal to surface.
125 %anttheta,antphi: angular poistions of surface patches to antenna position
126 %dA, dA_img: patch area & image patch areas
127 %dist: distance to patch from antenna
128 %patch_pos img_pos: vectorial positions of patches relative to antenna.
129 %khat= unit wavevector incident to each patch
130 [hemphi hemtheta antphi anttheta normal dA dA_img patch_pos img_pos...
131     distance khat]=build_lens(hemthetastep, hemphistep, Rsurf, ant_loc);
132
133 %Calculate the fields internal to the lens surface. Use data from the
134 %provided momentum file as well as geometry calculated above.
135 %antE & antH are the fields radiated from the antenna just inside the lens
136 %surface patches.
137 lambda=c/nlens/f; %[m] wavelength inside lens
138 [antE antH]=raytrace([pathin,fnamein],anttheta,antphi,patch_pos,...
139     lambda,nlens);
140
141 %Convert the internal fields to the basis wrt to the plane of incidence in
142 %preparation for refraction.
143 %TE_hat & TM_hat are unit vectors perpendicular to & within the plane of
144 %    incidence, both normal to k_hat

```

```

145 [TM_hat, TE_hat, Etransverse, Htransverse]=rect2transverse(normal, khat, ...
146         antE, antH);
147
148 %refractErect & refractHrect are electric and magnetic fields just outside
149 %the lens surface, in a rectangular basis
150 %The transmission coefficients are different depending on the style of
151 %AR-coating. The three cases are:
152 switch coatingflag
153     case {'no coating'}
154         [refractErect, refractHrect] =refraction_nocoat(Etransverse, ...
155             Htransverse, khat, TM_hat, TE_hat, normal, nlens, 1);
156         disp('Refraction through an uncoated surface');
157     case {'single layer'}
158         [refractErect, refractHrect] =refraction_1coat(Etransverse, ...
159             Htransverse, khat, TM_hat, TE_hat, normal, thickAR, nlens, nAR, 1, f);
160         disp('Refraction through a single layer AR-coating');
161     case {'three layers'}
162         [refractErect, refractHrect]=refraction_3coat(Etransverse, ...
163             Htransverse, khat, TM_hat, TE_hat, normal, nlens, n1, n2, n3, 1, thick1, ...
164             thick2, thick3, f);
165         disp('Refraction through a triple-layer AR-coating');
166 end
167
168 %Calculate surface currents (J,M) and their images (J_img, M_img).
169 [J, M, J_img, M_img]=surfacecurrents(refractErect, refractHrect, normal, Lext);
170 disp(['Finished after: ', num2str(toc), ' sec']);
171
172 %Do the Diffraction Integral to calculate far-fields (Efarfield, Hfarfield)
173 %at the angular positions (fftheta, ffphi)
174 [fftheta ffphi Efarfield Hfarfield]=Diffraction(ffthetastep, ffphistep, ...
175     J, M, J_img, M_img, dA, dA_img, patch_pos, img_pos, inc_img, Lext, f, ...
176     hemtheta, hemphi);
177
178 writefff(180/pi*ffphi, 180/pi*fftheta, Efarfield, [pathout, fnameout], ...
179     commentary);

```

### A.3 Construct Lens Geometry

```

1 function [hemphi hemtheta antphi anttheta normal dA dA_img patch_pos...
2     img_pos distance khat]=build_lens(hemthetastep, hemphistep, R, ant_loc)
3
4 %BUILD_LENS      build_lens(hem_theta_step, hem_phi_step, R, ant_loc)
5 %
6 %build_lens calculates lens geometry before any physics.
7 %This function accepts as arguments the angular step sizes for meshing the
8 %surface (hemthetastep & hemphistep) as well as lens radius (R) and the
9 %vectorial offset of the hemisphere's center from the antenna (ant_loc).
10 %It returns the folling in meshgrid format:
11 %HEMPHI & HEMTHETA: patch angular positions wrt center
12 %NORMAL: unit normal vectors to each patch
13 %dA: area if each patch, repeated for all 3-dimensions
14 %dA_img: area of image patches
15 %PATCH_POS: vector position of each patch wrt antenna
16 %img_pos: location of image patches

```

```

17 %DISTANCE: distance to each patch from antenna
18 %KHAT: unit angle of incidence
19 %ANTPHI & ANTTHEETA: patch angular positions wrt antenna
20 %All output vectors are in rectangular basis.
21 %Roger O'Brient          Oct 2009
22 %                        updated for image currents Jan 2010
23
24
25 %construct phi&theta coords on the hemisphere.  Exclude theta=0 (tip) and
26 %phi=360 (duplicate points).
27 [hemphi,hemtheta] = meshgrid(0:hemphistep:(2*pi-hemphistep),...
28     hemthetastep:hemthetastep:pi/2);
29 numtheta=size(hemtheta,1);
30 numphi=size(hemphi,2);
31
32 %unit normal
33 normal=cat(3,sin(hemtheta).*cos(hemphi),...
34     sin(hemtheta).*sin(hemphi),...
35     cos(hemtheta));
36
37 Lext=ant_loc(3);
38 %area of each patch & of image patches
39 dA=R^2.*sin(hemtheta);
40 dA= repmat(dA,[1 1 3]);
41 dA_img=dA;
42 %location of each patch
43 patch_pos=R*normal-...
44     repmat(permute(ant_loc,[3 2 1]),[numtheta,numphi,1]);
45
46 %img_pos = patch_pos(1:end-(Lext==0),:,:) ;
47 img_pos=patch_pos;
48 img_pos(:, :, 3)=-img_pos(:, :, 3);
49
50 distance=repmat(sqrt(dot(patch_pos,patch_pos,3)),[1 1 3]);
51 khat = patch_pos ./ distance ;
52
53 %atan is defined on [-90,90], but we need it over the full [0,360]
54 antphi=atan(khat(:, :, 2)./khat(:, :, 1))+...
55     pi*((khat(:, :, 1)<0)&(khat(:, :, 2)>0))+...
56     pi*((khat(:, :, 1)<0)&(khat(:, :, 2)<0))+...
57     2*pi*((khat(:, :, 1)>0)&(khat(:, :, 2)<0));
58 anttheta=acos((khat(:, :, 3)));

```

## A.4 Construct Fields just inside lens

```

1 function [antE antH]=raytrace(fff_file,anttheta,antphi,patch_pos,lambda,n)
2
3 %RAYTRACE raytrace(fff_file,anttheta,antphi,patch_pos,normal,lambda,khat,n)
4 %
5 %This fucntion calculates the internal fields of the lens just inside the
6 %surface.  It reads an ADS momentum generated file fff_file='*.fff' and
7 %interpolates electric field values at patches at (ANTTHEETA,ANTPHI).  It
8 %also accepts as input the rectangular location of each patch PATCH_POS,
9 %the wavelength in the material LAMBDA, and the wavespeed n to construct a

```

```

10 %propagator (greens function) that accounts for phase delays and 1/R field
11 %decay between the antenna and surface patches. It returns NxMx3 arrays of
12 %field antE and antH in rectangular coords. Note that the field must be
13 %the 1st argument of any cross product since Matlab conjugates the
14 %2nd argument.
15 %This function uses other functions NORMALIZE and SPHERE2RECT.
16 %
17 %Roger O'Brient      Jan 2010
18
19 eta=377; %[ohms]
20 %construct normalized wavevector
21 khat=cat(3,sin(anttheta).*cos(antphi),...
22         sin(anttheta).*sin(antphi),...
23         cos(anttheta));
24 ki=2*pi/lambda*khat;%[1/m] wavevector
25
26 %Read simulation file, interpolate field @ lens, convert to rect coords.
27 [simtheta,simphi,simEtheta,simEphi] = readfff(fff_file) ;
28 interpEphi = interp2(simphi,simtheta,simEphi,antphi,anttheta) ;
29 interpEtheta = interp2(simphi,simtheta,simEtheta,antphi,anttheta) ;
30 interpEsphere=cat(3,zeros(size(anttheta)),interpEtheta,interpEphi);
31 interpErect=sphere2rect(interpEsphere,anttheta,antphi);
32
33 %Construct propagator to account for phase delays to lens and 1/R field
34 %decay
35 R=repmat(sqrt(dot(patch_pos,patch_pos,3)), [1 1 3]);
36 propagator = repmat( exp( -j* dot( ki,patch_pos, 3 ) ), [ 1 1 3 ] ) ./R ;
37
38 %Now consctruct E and H fields at lens.
39 antE=interpErect.*propagator;
40 antH=-n*cross(antE,khat,3)/eta;

```

```

1 function [B]=normalize(A)
2 %Accepts an array of vectors (mxnx3), calculates the magnitude
3 %of each vector, and then divides that out to return an array of
4 %vectors that each have unit magnitude. The vectors can be in
5 %any basis.
6 %Roger O'Brient Oct 07
7 Amag=repmat(sqrt(dot(A,A,3)),[1 1 3]);
8 B=A./Amag;

```

## A.5 File reading & writing

```

1 function [theta phi Etheta Ephi] = readfff(fname)
2
3 fid=fopen( fname ) ;
4 if fid==-1
5     error(['error: cannot find file ',fname]);
6 end
7
8 data=[];
9
10 %fgetl reads a line, ignores new line character

```

```

11 line=fgetl(fid);
12
13
14 %feof=1 if at end of file, 0 otherwise.
15 while ~feof(fid)
16     %ignore if the line is empty or begins with '#', otherwise,
17     %concatinate with the data array as the next row.
18     if ~isempty( line ) && ~strcmp( line(1), '#' )
19         %convert line string into a vector of floating point variables
20         data = cat( 1, data, (sscanf( line, '%f' ))' );
21     end
22
23     %read next line before repeat
24     line=fgetl(fid);
25 end
26
27 %Data format is columns of:
28 %theta phi real(E_theta) imag(E_theta) real(E_phi) imag(E_phi)
29
30 %Now search the 1st two columns & ignore all the repeats
31 theta=unique(data(:,1))*pi/180 ;
32 phi=unique(data(:,2))*pi/180 ;
33
34 %Combine the real and imag components & put in meshgrid form
35 Etheta=reshape(complex(data(:,3),data(:,4)),...
36               length(theta),length(phi));
37 Ephi=reshape(complex(data(:,5),data(:,6)),...
38              length(theta),length(phi));
39 fclose(fid);

```

```

1 function wr=writefff(phi,theta,Efarfield,fname,commentary)
2 if nargin==5
3     commentary={};
4 end
5
6 %put the data in *.fff format so it could be fed back into ADS momentum
7 ffEtheta=Efarfield(:, :,2);
8 ffEphi=Efarfield(:, :,3);
9 dataout=permute(cat(3,theta,phi,...
10                  real(ffEtheta),imag(ffEtheta),...
11                  real(ffEphi),imag(ffEphi)),...
12                [3 1 2]);
13 fid=fopen(fname,'wt');
14 %first write comments, each line proceeded by a '#'
15 for i=1:length(commentary) ;
16     fprintf(fid,'# %s \n',commentary{i});
17 end
18
19 %Write data.  Sperate each theta cut with 'Begin Cut' & 'End Cut' as
20 %Momentum does.
21 for i=1:size(phi,2)
22     fprintf(fid,'# %s \n','Begin cut');
23     fprintf(fid,'%2.8f %2.8f %2.8e %2.8e %2.8e %2.8e \n', ...
24            dataout(:, :,i));
25     fprintf(fid,'# %s \n\n','End cut');
26 end
27 fclose(fid);

```

## A.6 Convert to & from spherical coordinates

```

1 function v_sphere=rect2sphere(v_rect,theta, phi)
2 %This converts an array of 3-d vectors in a rectangular basis to a
3 %spherical basis. All angles are in RADIANS. The theta and
4 %phi matrices should have been generated by meshgrid of
5 %dimensions mxn, while v_sphere should be mxn3. The program
6 %constructs |r>x|+|r>y|+|r>z| etc and dots this against the
7 %vectors in an xyz basis: V_r|r>(|r>x|+|r>y|+|r>z|)|v_rect>
8 %etc.
9 %Roger O'Brient Oct 07
10
11 rhat=cat(3, sin(theta).*cos(phi), sin(theta).*sin(phi)...
12     ,cos(theta)); %|r>x|+|r>y|+|r>z|
13 thetahat=cat(3, cos(theta).*cos(phi), cos(theta).*sin(phi)...
14     ,-sin(theta)); %|theta>x|+|theta>y|+|theta>z|
15 phihat=cat(3, -sin(phi), cos(phi),0*phi);
16     %|phi>x|+|phi>y|+|phi>z|
17
18 v_sphere=cat(3,...
19     dot(rhat,v_rect,3),... %V_r|r>(|r>x|+|r>y|+|r>z|)|v_rect>
20     dot(thetahat,v_rect,3),...%V_th|th>(|th>x|+|th>y|+|th>z|)|v_rect>
21     dot(phihat,v_rect,3)); %V_ph|ph>(|ph>x|+|ph>y|+|ph>z|)|v_rect>

```

```

1 function v_rect=sphere2rect(vsphere,theta, phi)
2 %This converts an array of 3-d vectors in a spherical basis to a
3 %rectangular basis. All angles are in RADIANS. The theta and
4 %phi matrices should have been generated by meshgrid of
5 %dimensions mxn, while v_sphere should be mxn3. The program
6 %constructs |x>r|+|x>th|+|x>ph| etc and dots this against the
7 %vectors in an xyz basis: V_x|x>(|x>r|+|x>th|+|x>ph|)|v_sph>
8 %etc.
9 %Roger O'Brient Oct 07
10
11 xhat=cat(3, sin(theta).*cos(phi), cos(theta).*cos(phi)...
12     ,-sin(phi)); %|x>r|+|x>th|+|x>ph|
13 yhat=cat(3, sin(theta).*sin(phi), cos(theta).*sin(phi)...
14     ,cos(phi)); %|y>r|+|y>th|+|y>ph|
15 zhat=cat(3, cos(theta), -sin(theta),zeros(size(phi)));
16     %|z>r|+|z>th|+|z>ph|
17
18 v_rect=cat(3,...
19     dot(xhat,vsphere,3),... %V_x|x>(|x>r|+|x>th|+|x>ph|)|v_sph>
20     dot(yhat,vsphere,3),...%V_y|y>(|y>r|+|y>th|+|y>ph|)|v_sph>
21     dot(zhat,vsphere,3)); %V_z|z>(|z>r|+|z>th|+|z>ph|)|v_sph>

```

## A.7 Convert to & from POI coordinates

```

1 function [TM_hat,TE_hat,Eplane,Hplane]=rect2transverse(normal,k_hat,...
2     Erect,Hrect)
3 % This function accepts unit normal (NORMAL) and unit incident

```



```

4 %           (KHAT) vectors and constructs a basis perpendicular (TE_HAT)
5 %           and within the plane of incidence (TM_HAT). The basis is
6 %           TM_HAT=TE_HAT X KHAT
7 %
8 %           TE stands for Transverse-Electric component while TM for
9 %           Transverse-Magnetic component.
10 %          It then resolves the provided E and H vectors into components
11 %          parallel (E_TM) and perpendicular (E_TE), where the output
12 %          format is an mxn x 3 matrix. The mxn columns refer to specific
13 %          angles of the lens surface patches and the components on the
14 %          3-element dimension are:
15 %              1. along K_HAT (which will be zero)
16 %              2. along TM_HAT
17 %              3. along TE_HAT
18 %          The function returns the two new basis vectors as well as the
19 %          fields in that basis.
20 %          This function calls the custom function NORMALIZE which forces a
21 %          matrix of 3-vectors to be normal.
22 %          WARNING: Matlab's native "dot" function takes the conjugate of
23 %          it's first argument, and "cross" conjugates the second!
24 %          Roger O'Brient Oct 07
25
26 %set up the new basis.
27 normal=normalize(normal);
28 k_hat=normalize(k_hat);
29 TE_hat=cross(normal,k_hat,3);
30 TE_hat=normalize(TE_hat);
31 TM_hat=cross(TE_hat,k_hat,3);
32 TM_hat=normalize(TM_hat);
33
34 %Find components in new basis. Basis vectors must be the first argument
35 %since they are real and Matlab will automatically conjugate those.
36 E_k=dot(k_hat,E_rect,3);
37 E_tm=dot(TM_hat,E_rect,3);
38 E_te=dot(TE_hat,E_rect,3);
39
40 H_k=dot(k_hat,H_rect,3);
41 H_te=dot(TM_hat,H_rect,3);
42 H_tm=dot(TE_hat,H_rect,3);
43
44 %construct vectors in new basis as described in header
45 Eplane=cat(3,E_k,E_tm,E_te);
46 Hplane=cat(3,H_k,H_te,H_tm);

```

```

1 function [Erect,Hrect]=transverse2rect(normal,k_hat,...
2           Etransverse,Htransverse)
3 %This function accepts unit normal (NORMAL) and unit incident
4 % (KHAT) vectors and constructs a basis perpendicular (TE_HAT)
5 %and within the plane of incidence (TM_HAT). These names are in reference
6 %to the orientation of the electric fields of those components; e.g the
7 %electric field of the TM-component resides in the plane of incidence.
8 %The basis is
9 %           TM_HAT=TE_HAT x K_HAT
10 %and all vectors are ordered:
11 %           1.k_hat
12 %           2.TM_hat
13 %           3.TE_hat

```

```

14 %It then resolves the E and H vectors provided by the user in
15 %that basis back into a rectangular basis, where the output
16 %format is an mxn3 matrix. The mxn columns refer to specific
17 %angular positions of the lens surface patches and the components on the
18 %3-element dimension are:
19 %     1. x_hat
20 %     2. y_hat
21 %     3. z_hat
22 %This function calls the custom function NORMALIZE which forces a
23 %matrix of 3-vectors each be of unit length.
24 %WARNING: Matlab's native "dot" function takes the conjugate of
25 %it's first argument, and "cross" conjugates the second!
26 %Roger O'Brient Oct 2009
27
28 %Set up the new basis. These are in cartesian coordinates.
29 normal=normalize(normal);
30 k_hat=normalize(k_hat);
31 TE_hat=cross(normal,k_hat,3);
32 TE_hat=normalize(TE_hat);
33 TM_hat=cross(TE_hat,k_hat,3);
34 TM_hat=normalize(TM_hat);
35
36 %extract components in the transverse-basis.
37 E_k=repmat(Etransverse(:,:,1),[1 1 3]);
38 E_tm=repmat(Etransverse(:,:,2),[1 1 3]);
39 E_te=repmat(Etransverse(:,:,3),[1 1 3]);
40
41 H_k=repmat(Htransverse(:,:,1),[1 1 3]);
42 H_te=repmat(Htransverse(:,:,2),[1 1 3]);
43 H_tm=repmat(Htransverse(:,:,3),[1 1 3]);
44
45 %Assemble vectors in rectangular coordinates.
46 Erect=E_k.*k_hat+E_tm.*TM_hat+E_te.*TE_hat;
47 Hrect=H_k.*k_hat+H_te.*TM_hat+H_tm.*TE_hat;

```

## A.8 Refraction

There are three options: no coating, one-layer coating, and three layers. For brevity's sake (too late...), I have excluded the one-layer function which is similar to the three-layer.

```

1 function [transErect,transHrect] =refraction_nocoat (Etransverse,Hplane,...
2     k_hat, TM_hat, TE_hat, normal, nlens, noutside)
3 %REFRACTION_NOCOAT refraction_nocoat (Etransverse,Hplane,...
4 %     k_hat, TE_hat, TM_hat, normal, nlens, noutside)
5 %This function refracts the incident fields (EPLANE,HPLANE) in
6 %Plane-Of-Incidence (POI) coords into refracted fields
7 % (TRANSERECT,TRANSHRECT), rectangular coords. All are mxn3 arrays.
8 %It accepts the POI basis vectors (KHAT, TM_hat, TE_hat) as well as the
9 %surface normals NORMAL and the indices inside (N1) and outside (N2) the
10 %lens.
11 %This function invokes Snell's law to calculate the new unit wavevector
12 %khatprime. It calculates Fresnel Coefficients, and then the transmitted
13 %fields. The TE component of the E-fields remains in the same
14 %position, as does the TM component of the H-fields. The others rotate

```

```

15 %about the basis vectors TE_hat.  HTRANS is calculated from ETRANS and
16 %khatprime.
17 %Roger O'Brient Jan 2010
18
19 eta=377; %[Ohms] Impedance Free Space
20
21 %Snell's law rotates pointing vector in khat-TM_hat plane by
22 %angleDelta:
23 angle_inc=acos(dot(k_hat,normal,3)); %[radians]
24 angle_trans=asin(nlens/noutside*sin(angle_inc)); %[radians]
25 angleDelta=angle_trans-angle_inc; %[radians]
26 %The wavevector for the transmitted ray is just a rotation of the incident
27 %wavevector by angle_Delta about the TE_hat axis. transk_hat will be part
28 %of a new basis-set for the fields.
29 transk_hat= repmat(cos(angleDelta),[1 1 3]).*k_hat+...
30     repmat(sin(angleDelta),[1 1 3]).*TM_hat;
31
32 %Fresnel Coefficients (unitless):
33 R_te=(nlens*cos(angle_inc)-noutside*cos(angle_trans))./...
34     (nlens*cos(angle_inc)+noutside*cos(angle_trans));
35 T_te=1+R_te;
36 R_tm=(-noutside*cos(angle_inc)+nlens*cos(angle_trans))./...
37     (noutside*cos(angle_inc)+nlens*cos(angle_trans));
38 T_tm=nlens/noutside*(1-R_tm);
39
40 %Transmitted E-field perpendicular to POI (Etransverse(:, :, 3)) points in
41 %same direction before and after refraction. The E-field within the POI
42 %Etransverse(:, :, 2) rotates like khat by angleDelta
43
44 transEplane=cat(3,-T_tm.*Etransverse(:, :, 2).*sin(angleDelta),...
45     T_tm.*Etransverse(:, :, 2).*cos(angleDelta),...
46     T_te.*Etransverse(:, :, 3));
47
48 %convert to rectangular coords
49 [transErect,transHrect]=transverse2rect(normal,k_hat,transEplane,...
50     transEplane);
51
52 %Construct Magnetic-field
53 transHrect=1/eta*cross(transk_hat,transErect,3);

```

```

1 function [transErect,transHrect] =refraction_3coat(Etransverse,...
2     Htransverse,k_hat, TM_hat, TE_hat, normal,nlens,n1,n2,n3,noutside,...
3     thick1,thick2,thick3,f)
4 %REFRACTION_3COAT refraction_3coat(Etransverse,Htransverse,k_hat,TE_hat,...
5 %TM_hat,normal,nlens,n1,n2,n3,noutside,thick1,thick2,thick3,f)
6 %This function refracts the incident fields (ETRANSVERSE,HTRANSVERSE) in
7 %Plane-Of-Incidence (POI) coords into refracted fields
8 %(TRANSERECT,TRANSHRECT). All are mxnx3 arrays. It accepts
9 %the basis vectors (KHAT, TM_HAT, TE_HAT) as well as the surface
10 %normals NORMAL and the indices inside (NLEN) and outside (NOUTSIDE) the
11 %lens. It also accepts the thicknesses of the 3 layers THICK1-THICK3 and
12 %their indices N1-N3. Layer 1 is the inner-most, Layer 3 the outer-most.
13 %This function invokes Snell's law to calculate the new unit wavevector
14 %khatprime, which is just a rotation about the TE_hat basis.
15 %It calculates Fresnel Coefficients for a three layer AR-coating by
16 %forcing the fields to be continuous at the boundaries between the media,
17 %and then calculates transmitted fields. The TE component of the E-fields

```

```

18 %remains in the same position, as does the TM component of the H-fields.
19 %The others rotate about the basis vectors TE_hat just like the wavevector.
20 %TRANSHRECT is calculated from TRANSERECT and khatprime.
21 %This function uses the function matrix_mult4d, since matlab cannot
22 %natively do matrix multiplication on arrays of rank>2.
23 %Roger O'Brient          Jan 2010
24
25 c=3*10^8; %[m/s] Speed light free space
26 eta=377; %[Ohms] Impedance Free Space
27
28 lambda_o=c/f; %[m] wavelength outside lens
29 ko=(2*pi/lambda_o); %[1/m] wavenumber outside lens
30
31 %Snell's law rotates pointing vector in khat-TM_hat plane at every coating
32 %interface. The net rotation is angleDelta, which is as if there never was
33 %no intermediate layers. But we need the intermediate angles for
34 %calculating the transmission coefficients.
35 angle_inc=acos(dot(k_hat,normal,3)); %[radians]
36
37 angle1=asin(nlens/n1*sin(angle_inc)); %[radians]
38 angle2=asin(n1/n2*sin(angle1)); %[radians]
39 angle3=asin(n2/n3*sin(angle2)); %[radians]
40
41 angle_trans=asin(nlens/noutside*sin(angle_inc)); %[radians]
42 angleDelta=angle_trans-angle_inc; %[radians]
43 %The wavevector for the transmitted ray is just a rotation of the incident
44 %wavevector by angle_Delta about the TE_hat axis. transk_hat can be part
45 %of a new basis-set for the fields, although I keep all vectors in the
46 %original incident basis
47 transk_hat=repmat(cos(angleDelta),[1 1 3]).*k_hat+...
48     repmat(sin(angleDelta),[1 1 3]).*TM_hat;
49
50 %Construct Transmission coefficients using the well known algorithm
51 %discussed in section 9.7.1 of Hect's Optics using transfer matrices.
52 %Construct transfer matrices: [E1;H1]=M*[E2;H2] for each layer and then
53 %multiply them
54
55 Y_lens=nlens/eta*cos(angle_inc);
56
57 Y_1_te=n1/eta*cos(angle1);
58 Y_1_tm=n1/eta*sec(angle1);
59 Y_2_te=n2/eta*cos(angle2);
60 Y_2_tm=n2/eta*sec(angle2);
61 Y_3_te=n3/eta*cos(angle3);
62 Y_3_tm=n3/eta*sec(angle3);
63
64 Y_out=noutside/eta*cos(angle_trans);
65
66 h1=n1*thick1*cos(angle1);
67 h2=n2*thick1*cos(angle2);
68 h3=n3*thick3*cos(angle3);
69
70 M1_te=cat(3,cat(4,cos(ko*h1),          i*sin(ko*h1)./Y_1_te),...
71     cat(4,i*sin(ko*h1).*Y_1_te,      cos(ko*h1)));
72 M1_tm=cat(3,cat(4,cos(ko*h1),          i*sin(ko*h1)./Y_1_tm),...
73     cat(4,i*sin(ko*h1).*Y_1_tm,      cos(ko*h1)));
74
75 M2_te=cat(3,cat(4,cos(ko*h2),          i*sin(ko*h2)./Y_2_te),...
76     cat(4,i*sin(ko*h2).*Y_2_te,      cos(ko*h2)));

```

```

77 M2_tm=cat(3,cat(4,cos(ko*h2),          i*sin(ko*h2)./Y_2_tm),...
78           cat(4,i*sin(ko*h2).*Y_2_tm,  cos(ko*h2)));
79
80 M3_te=cat(3,cat(4,cos(ko*h3),          i*sin(ko*h3)./Y_3_te),...
81           cat(4,i*sin(ko*h3).*Y_3_te,  cos(ko*h3)));
82 M3_tm=cat(3,cat(4,cos(ko*h3),          i*sin(ko*h3)./Y_3_tm),...
83           cat(4,i*sin(ko*h3).*Y_3_tm,  cos(ko*h3)));
84
85 %Multiply the transfer matrices.
86 Mtot_te=matrix_mult4d(M1_te,matrix_mult4d(M2_te,M3_te));
87 Mtot_tm=matrix_mult4d(M1_tm,matrix_mult4d(M2_tm,M3_tm));
88
89 %Construct transmission coefficients for the E-fields
90 T_te=2*Y_out./(Y_out.*Mtot_te(:, :, 1, 1)+...
91              Y_out.*Y_lens.*Mtot_te(:, :, 1, 2)+...
92              Mtot_te(:, :, 2, 1)+...
93              Y_lens.*Mtot_te(:, :, 2, 2));
94
95 T_tm=2*Y_out./(Y_out.*Mtot_tm(:, :, 1, 1)+...
96              Y_out.*Y_lens.*Mtot_tm(:, :, 1, 2)+...
97              Mtot_tm(:, :, 2, 1)+...
98              Y_lens.*Mtot_tm(:, :, 2, 2));
99
100 %Transmitted E-field perpendicular to POI (Etransverse(:, :, 3)) points in
101 %same direction before and after refraction. The E-field within the POI
102 %(Etransverse(:, :, 2))rotates by angleDelta like khat.
103
104 transEplane=cat(3,-T_tm.*Etransverse(:, :, 2).*sin(angleDelta),...
105                T_tm.*Etransverse(:, :, 2).*cos(angleDelta),...
106                T_te.*Etransverse(:, :, 3));
107
108 %covert to rectangular coords
109 [transErect,transHrect]=transverse2rect(normal,k_hat,transEplane, ...
110                                       transEplane);
111
112 %Construct Magnetic-field
113 transHrect=1/eta*cross(transk_hat,transErect,3);

```

```

1 function [C]=matrix_mult4d(A,B)
2 %4DMATRIX_MULT C=matrix_mult4d(A,B)
3 %This function multiplies two mxnx2x2 arrays A and B in the last two
4 %indices according to standard matrix multiplication definition. It
5 %returns a mxnx2x2 array where each of the mth,nth 2x2 array is the matrix
6 %product of the corresponding ones from A and B. Matlab
7 %does not have a native way of doing this, but this is needed for 2x2
8 %multiplication in the AR-coating algorithm in our ray-tracing code.
9 %
10 %Roger O'Brient      Feb 2010
11
12 C11=A(:, :, 1, 1).*B(:, :, 1, 1)+A(:, :, 1, 2).*B(:, :, 2, 1);
13 C12=A(:, :, 1, 1).*B(:, :, 1, 2)+A(:, :, 1, 2).*B(:, :, 2, 2);
14 C21=A(:, :, 2, 1).*B(:, :, 1, 1)+A(:, :, 2, 2).*B(:, :, 2, 1);
15 C22=A(:, :, 2, 1).*B(:, :, 1, 2)+A(:, :, 2, 2).*B(:, :, 2, 2);
16
17 Ccol1=cat(4,C11,C12);
18 Ccol2=cat(4,C21,C22);
19 C=cat(3,Ccol1,Ccol2);

```

## A.9 diffraction

```
1 function [J,M,J_img,M_img]=surfacecurrents(E,H,normal,Lext)
2 %       Function takes as input mxn3 arrays of 3-D vectors in
3 %       rectangular basis corresponding to Electric Fields E, Magnetic
4 %       Fields H, and unit normal vectors NORMAL at the lens surface.
5 %       It constructs fictitious electric J and magnetic currents M on
6 %       the surface to assist with the Huygens Integral. These vectors
7 %       are also in a rectangular coordinate basis. It also returns
8 %       the image currents J_IMG & M_IMG, but reflects them to ensure
9 %       that the electric fields at the ground plane are normal and
10 %      magnetic fields are tangential.
11 %      Roger O'Brient           October 2009
12 %                               updated for image currents Jan 2010
13
14 J= cross(normal,H,3);
15 M= -cross(normal,E,3);
16
17 J_img=J;
18 J_img(:,:,1:2)=-J_img(:,:,1:2);
19
20 M_img=M;
21 M_img(:,:,3)=-M_img(:,:,3) ;
```

```
1 function [fftheta ffphi Efarfield Hfarfield]=Diffraction(ffthetastep,...
2     ffphistep,J,M,J_img,M_img,dA,dA_img,patch_pos,img_pos,inc_img,Lext,...
3     f,hemtheta,hemphi)
4 %DIFFRACTION
5 %This function returns the far-fields (EFARFIELD,HFARFIELD) at angular
6 %positions (FFTHETA,FFPHI). It's arguments are theta and phi steps for
7 %the far field, surface currents and their images, the locations of those
8 %currents, patch surface areas, a flag to include the images, and the
9 %extension length LEXT.
10 %The code uses a common Fourier Transform algorithm outline in most
11 %antenna textbooks for radiation through an aperture. This was developed
12 %on a laptop with 0.5Gb RAM, which was insufficient to store 5-index arrays
13 %needed for the diffraction calculations. So it uses a loop instead and it
14 %reports progress to the user with the function PROGMETER. If all angles
15 %steps are 3deg, then the code takes 30-40sec to execute on my laptop with
16 %1.8GHz processor.
17 %Roger O'Brient   Aug2009
18
19 c=3*10^8; %[m/s]
20 eta=377; %[Ohms]
21 if Lext==0
22     remove=1;
23 else
24     remove=0;
25 end
26 tic ;
27 [ffphi,fftheta]=meshgrid(0:ffphistep:(2*pi),0:ffthetastep:(pi/2));
28 Efarfield=zeros(cat(2,size(ffphi),3));
29 Hfarfield=zeros(cat(2,size(ffphi),3));
30 progmeter(0, 'Performing Diffraction Integral')
31
```

```

32 for p=1:size(ffphi,2)
33     for t=1:size(fftheta,1)
34         %Construct the wavevector of the total wave at the far-field
35         %angular position (fftheta,ffphi
36         theta=fftheta(t,p);
37         phi=ffphi(t,p);
38         ffk=2*pi*f/c*cat(3,...
39         repmat(sin(theta).*cos(phi),[size(hemtheta,1) size(hemphi,2) 1]),...
40         repmat(sin(theta).*sin(phi),[size(hemtheta,1) size(hemphi,2) 1]),...
41         repmat(cos(theta),[size(hemtheta,1) size(hemphi,2) 1]));
42         %Construct Propagator to far field. This only accounts for phase
43         %difference since intensity decay is roughly the same at all
44         %points. The Propagator for the images is almost the same, except
45         %excludes equator points if Lext=0.
46         Propagator = repmat(exp(i*sum(ffk.*patch_pos,3)),[1 1 3]) ;
47         if inc_img
48             Propagator_img=repmat(exp(i*sum(ffk(1:end-remove,,:)).*...
49             img_pos,3)),[1 1 3]);
50         end
51         %Construct Far-field Magnetic (N) and Electric (L) Vector
52         %potentials. These exclude common factors of phase delay and 1/r
53         %field decay.
54         Nrect=sum(sum(J.*Propagator.*dA,1),2)+...
55             inc_img*sum(sum(J_img.*Propagator_img.*dA_img,1),2);
56         Lrect=sum(sum(M.*Propagator.*dA,1),2)+...
57             inc_img*sum(sum(M_img.*Propagator_img.*dA_img,1),2);
58         N=rect2sphere(Nrect,theta,phi);
59         L=rect2sphere(Lrect,theta,phi);
60
61         %Fields are derivatives of the potentials:
62         Efarfield(t,p,2)=- (L(:, :, 3)+eta*N(:, :, 2));
63         Efarfield(t,p,3) = (L(:, :, 2)-eta*N(:, :, 3));
64         Hfarfield(t,p,2) = (N(:, :, 3)-L(:, :, 2)/eta);
65         Hfarfield(t,p,3)=- (N(:, :, 2)+L(:, :, 3)/eta);
66     end
67     progmeter(phi/(2*pi));
68 end
69
70 %normalize beams to peak
71 EPower=dot(Efarfield,Efarfield,3);
72 HPower=dot(Hfarfield,Hfarfield,3);
73 Efarfield=Efarfield/sqrt(max(max(EPower)));
74 Hfarfield=Hfarfield/sqrt(max(max(HPower)));
75 % ffEtheta=Efarfield(:, :, 2);
76 % ffEphi=Efarfield(:, :, 3);
77 % ffHtheta=Hfarfield(:, :, 2);
78 % ffHphi=Hfarfield(:, :, 3);
79 progmeter done
80 disp(['Diffraction Calculation finished after ',num2str(toc),' sec']);

```